

# AN ESSENTIALLY STATISTICAL APPROACH TO THERMODYNAMIC PROBLEM—II

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## INTRODUCTION

In a recent paper \* (1953) the author has stressed the necessity of an essentially statistical approach to the thermodynamic problems and has developed a general statistical method for investigating thermodynamic behaviour of a system. There, for simplicity, the system has been assumed to be composed of matter of (chemically) single type, and the contribution of radiation energy, if any, in the behaviour of the system has been taken to be negligibly insignificant. In the present paper, the proposed statistical method will be shown to be equally suitable for investigations of thermodynamic behaviour of system of different types, viz. those of black-body radiations, those of free radiation and matter of single chemical types, and those of matter of chemically different types. In agreement with other usual theories and with the assumptions of the paper I, it is assumed that the amount of energy and matter can only be discrete. In every case, the usual formulae, as given by Fowler (1936), Tolman (1946) and others, have been obtained.

## PART I. AN ASSEMBLY OF FREE RADIATION

### 1. *Formulation of Problem: The Law of Probability-distribution*

Here, the system means a definite amount of energy  $E$ , in the form of free radiation in a definite volume in a definite environment. As in the paper I, the system is to be taken as open, i.e. there is no hindrance of energy-exchange between the system and the environment. Now, since there is no question of matter-exchange in this case with the environment, so in consistency with the trend of discussions made in the paper I, only one a priori probability  $z$  for each unit of energy to be in the volume under consideration will be introduced. Then, the probability of occurrences of a system in a state specified by  $E$  is given by

$$P(z, E) = \frac{W(E)z^E}{\sum_{E=0}^{\infty} W(E)z^E} = \frac{W(E)z^E}{f(z)}, \quad \dots \dots \dots (1)$$

where 
$$f(z) = \sum_{E=0}^{\infty} W(E)z^E \quad \dots \dots \dots (2)$$

The above may be looked upon as a form of Baye's theorem. As before, for determining  $z$  in agreement with the usual ideas of statistical equilibrium, the 'real' value  $E_0$  (taken to be known in the sense of classical thermodynamics) corresponds

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\* This paper will be referred to as the paper I in the present discussion.



Thus,  $\Phi(z)$ , i.e.  $P(z, E)$ , has only one maximum. Therefore  $z_0$  is the unique value of  $z$  corresponding to the stationary state.

4. *Influence of Variations of Environments*

Now as in the paper I, if the variations of parameters, involved in the definitions of the environments, be taken into account, then the equation of continuity of energy after usual averaging can be written as

$$\overline{dE} = \sum_i \frac{\partial \overline{E}}{\partial x_i} dx_i + d'Q, \quad \dots \quad (8)$$

where

$$\left(-\frac{\partial \overline{E}}{\partial x_i}\right) = \frac{\sum_{E=0}^{\infty} \left(-\frac{\partial E}{\partial x_i}\right) W(E) z_0^E}{\sum_{E=0}^{\infty} W(E) z_0^E} = -\frac{1}{\left(\log \frac{1}{z_0}\right)} \cdot \frac{\partial}{\partial x_i} \{\log f(z_0)\} \quad \dots \quad (9)$$

Now

$$\begin{aligned} \left(\log \frac{1}{z_0}\right) d'Q &= \left(\log \frac{1}{z_0}\right) \left\{ \overline{dE} - \sum_i \frac{\partial \overline{E}}{\partial x_i} dx_i \right\} \\ &= d \left\{ \left(\log \frac{1}{z_0}\right) \overline{E} \right\} + \sum_i \frac{\partial}{\partial x_i} \{\log f(z_0)\} dx_i + \frac{\partial}{\partial z_0} \{\log f(z_0)\} \\ &= d \left\{ \left(\log \frac{1}{z_0}\right) \overline{E} + \log f(z_0) \right\} \quad \dots \quad (10) \end{aligned}$$

Then, on putting

$$z_0 = e^{-\frac{1}{kT}}, \quad \dots \quad (11)$$

after choosing the scale properly as in the paper I, one has

$$dS = d \left\{ \frac{E}{T} + k \log F(T) \right\}. \quad \dots \quad (12)$$

5. *Laws of Microscopic Distributions*

Now, here if the radiation be assumed to be composed of the photon and if number of ways in which a particular  $n_\nu$  (number of photons) with energy  $\epsilon_\nu$  be realised when the total energy of the system  $E$  is denoted by  $w'(E - n_\nu \epsilon_\nu)$ , then the expected value of  $\bar{n}_\nu$  is given by

$$\begin{aligned} \bar{n}_\nu &= \frac{\sum_{E=0}^{\infty} \sum_{n_\nu=0}^{\infty} n_\nu w'(E - n_\nu \epsilon_\nu) z_0^E}{\sum_{E=0}^{\infty} W(E) z_0^E} \\ &= \frac{\left( \sum_{n_\nu=0}^{\infty} n_\nu z_0^{n_\nu \epsilon_\nu} \right) \left[ 1 + \sum_E w'(E) z_0^E \right]}{\left( \sum_{n_\nu=0}^{\infty} z_0^{n_\nu \epsilon_\nu} \right) \left[ 1 + \sum_E w'(E) z_0^E \right]} \end{aligned}$$

$$= z_0^{\epsilon_\nu} \frac{\partial}{\partial(z_0^{\epsilon_\nu})} \left\{ \log \left( 1 + \sum_{n_\nu}^{\infty} z_0^{n_\nu \epsilon_\nu} \right) \right\} \dots \dots \dots (13)$$

Then, the spectral-distribution of energy becomes

$$\bar{n}_\nu = \frac{A_\nu z_0^{\epsilon_\nu}}{1 - z_0^{\epsilon_\nu}} = \frac{A_\nu}{e^{\epsilon_\nu/kT}} \dots \dots \dots (14)$$

This is the well-known Planck's law.

*Remarks*

Now, the assumption of existence of  $W(E)$ , the total number of realizations in which  $E$  can be realized in different ways in different frequencies, implies the exchange of energies amongst different frequencies, i.e. the existence of some matter in contact with the system consisting of radiation under consideration has been tacitly assumed. So, the present case is the case of black-body radiation.

PART II. AN ASSEMBLY CONSISTING OF RADIATION AND MATTER

1. *Description of Assembly*

Here, as in the paper I, a system in a definite thermodynamic state will mean a definite amount of energy and a definite amount of matter observed (in the sense of classical thermodynamics) in a definite volume in a definite environment. Here, as in the paper I, the system will be taken to be open with respect to energy and matter. Of course, for simplicity at present, the system will be assumed to be composed of (chemically) single type of material.

Now, as before, a constant a priori probability for a unit quantity of energy to exist in the volume and another constant a priori probability for a unit quantity of matter to exist in the volume will be introduced. Then, the probability of occurrence of the system in a state specified by  $E$  and  $M$  is given by

$$P(t, z, M, E) = \frac{W(E, M)t^M z^E}{\sum_E \sum_M W(E, M)t^M z^E} = \frac{W(E, M)t^M z^E}{f(t, z)}, \dots \dots (15)$$

where

$$f(t, z) = \sum_E \sum_M W(E, M)t^M z^E. \dots \dots \dots (16)$$

This expression is same as that in the previous paper (Dutta, 1953). The only difference is that here  $W(E, M)$  can always be written as

$$\sum_{E_1 + E_2 = E} W_1(E_1) W_2(E_2, M) = W(E, M), \dots \dots (17)$$

where  $E_1$  is the amount of energy in the form of free radiation and  $E_2$  that as the energy of matter. Obviously  $W_1(E_1)$  is the function used in the part I of this paper and  $W_2(E_2)$  is that used in the paper I. Then, from the theorem, stated and proved in the paper I, one gets

$$f(t, z) = f_1(z)f_2(t, z). \dots \dots \dots (18)$$

As before, the distribution-parameters  $t, z$  corresponding to the real values of  $E_0, M_0$  (taken to be known from the observations in classical sense) are to be

specified from the principle that the real values of  $E_0$  and  $M_0$  correspond to the maximum of the probability of occurrence considered as function of  $t$  and  $z$ .

*Specification and Uniqueness of  $t_0$  and  $z_0$ , etc.*

In this case, equations specifying values of  $t_0$  and  $z_0$  and calculations showing the uniqueness of specified values and discussions of the effects of variations in the environments are same in the previous paper and so are not repeated here.

*Laws of Microscopic Distributions*

The previous discussions are mainly of microscopic nature and have been put forward without any specific assumptions about the microscopic nature of the system. Here, matter will be assumed to be composed of molecules and radiation of photons. As stated in the paper I,  $E_1$  and  $E_2$  are both to be taken as discrete. Let us write

$$w(E, N) = W(E, M) = w_1(E_1)w_2(E_2, N). \quad \dots \quad (19)$$

Let  $\epsilon_\nu$  be the energy of one of the micro-energy state of photons and  $n_\nu$  be the number of photons with energy  $\epsilon_\nu$  at any instant. Then, on writing

$$t'_0 = t_0^m, \quad \dots \quad (20)$$

one gets

$$\begin{aligned} n_\nu &= \frac{\sum_{E=0}^{\infty} \sum_{N=0}^{\infty} \sum_{n_\nu=0}^{\infty} n_\nu w'_1(E_1 - n_\nu \epsilon_\nu) w_2(E_2, N) t_0^N z_0^E}{\sum_{\substack{E=0 \\ E_1+E_2=0}}^{\infty} \sum_{N=0}^{\infty} w'_1(E_1) w'_2(E_2, N) t_0^N z_0^E} \\ &= \frac{\left( \sum_{n_\nu=0}^{\infty} n_\nu z_0^{n_\nu \epsilon_\nu} \right) \left[ 1 + \sum w'_1(E_1) w'_2(E_2, N) t_0^N z_0^E \right]}{\left( 1 + \sum_{n_\nu=1}^{\infty} z_0^{n_\nu \epsilon_\nu} \right) \left[ 1 + \sum \sum w'_1(E_1) w'_2(E_2, N) t_0^N z_0^E \right]} \\ &= z_0^{\epsilon_\nu} \frac{\partial}{\partial z_0^{\epsilon_\nu}} \left\{ \log \left( 1 + \sum_{n_\nu=0}^{\infty} z_0^{n_\nu \epsilon_\nu} \right) \right\}, \\ t'_0 &= t_0^m \quad \dots \quad (21) \end{aligned}$$

where  $w'_1(E)$  is the number of partition of  $E_1$ , into sets of  $\epsilon_0, \epsilon_1, \dots, \epsilon_\nu, \dots$  of which no part is equal to  $\epsilon_\nu$ . Then the spectral-distribution of radiation energy becomes

$$\bar{n}_\nu = \frac{A_\nu z_0^{\epsilon_\nu}}{1 - z_0^{\epsilon_\nu}} = \frac{A_\nu}{e^{\epsilon_\nu/kT} - 1} \quad \dots \quad (22)$$

Similarly, the microscopic distribution for molecules with respect of energy can be obtained. This is again same as that in the paper I.

Formula for fluctuations can be obtained similarly and similar to these in the previous paper.

PART III. ASSEMBLY COMPOSED OF MATTER OF TWO DIFFERENT TYPES

1. *Description of System: Law of Probability-distribution*

The system means certain amount of  $E$  and certain amounts  $M_1$  and  $M_2$  of matters of two (chemically) different types, contained in a definite volume in a definite environment. For simplicity, it will be assumed that there is no chemical reaction in the assembly and that the condition of the system and the environment is such that the contribution in the energy due to the free radiation in the system may be taken as insignificantly small and so will be neglected.

Here, as before, a constant a priori  $z$  for each unit of energy to occur within the volume under consideration and constant a priori probabilities  $t_1$  and  $t_2$  for unit masses of matter of types (1) and (2) to be in the volume respectively will be introduced. Then, the probability of occurrences of the system to be in a state specified by  $E, M_1$  and  $M_2$  is given by

$$\begin{aligned}
 P(t_1, t_2, z, E, M_1, M_2) &= \frac{W(E, M_1, M_2)z^E t_1^{M_1} t_2^{M_2}}{\Sigma \Sigma \Sigma W(E, M_1, M_2)z^E t_1^{M_1} t_2^{M_2}} \\
 &= \frac{W(E, M_1, M_2)z^E t_1^{M_1} t_2^{M_2}}{f(t_1, t_2, z)}, \quad \dots \quad \dots \quad (23)
 \end{aligned}$$

where

$$f(t_1, t_2, z) = \sum_E \sum_{M_1} \sum_{M_2} W(E, M_1, M_2)z^E t_1^{M_1} t_2^{M_2}. \quad \dots \quad \dots \quad (24)$$

The above may be looked upon as same form of Baye's theorem. As before, for determining  $z, t_1, t_2$  for the distribution law, in agreement with usual ideas of statistical equilibrium, the real values of  $E_0, M_{10}, M_{20}$  (taken to be known from the observations in the classical sense) are assumed to correspond to the maximum of the probability of occurrence considered as functions of  $z, t_1, t_2$  for fixed values of  $E = E_0, M_1 = M_{10}, M_2 = M_{20}$ . Thus, the estimation of parameters  $z, t_1$  and  $t_2$  is similar to that in statistics by principle of maximum likelihood.

2. *Behaviour of Functions of  $W(E, M_1, M_2), f(t_1, t_2, z), etc.$*

All these functions are similar to those in the paper I, and in the parts I and II of this paper, and some similar properties. As in the paper I and the part II of this paper, it can be easily shown that

$$f(t_1, t_2, z) = f_1(t_1, z)f_2(t_2, z), \quad \dots \quad \dots \quad \dots \quad (25)$$

where

$$W(E, M_1, M_2) = \sum_{E_1 + E_2 = E} W_1(E_1, M_1)W_2(E_2, M_2), \quad \dots \quad \dots \quad (26)$$

$$f_1(t_1, z) = \sum_{E_1} \sum_{M_1} W_1(M_1, E_1)z^{E_1} t_1^{M_1}, \quad \dots \quad \dots \quad (27)$$

$$f_2(t_2, z) = \sum_{E_2} \sum_{M_2} W_2(M_2, E_2)z^{E_2} t_2^{M_2}, \quad \dots \quad \dots \quad (28)$$

and

$$E = E_1 + E_2, \quad \dots \quad \dots \quad \dots \quad (29)$$

$E$  being the total energy of the system and  $E_1, E_2$  being energies shared by matters of types (1) and (2).

As before,

$$P(t_1, t_2, z) \geq 0, \quad \dots \dots \dots (30)$$

and for infinite rarefactions,  $t_1 = 0, t_2 = 0, z = 0,$

$$P(t_1, t_2, z) = 0, \quad \dots \dots \dots (31)$$

and for infinite condensation,  $t_1 = 1, t_2 = 1, z = 1,$

$$P(t_1, t_2, z) = 0. \quad \dots \dots \dots (32)$$

Then,  $P(t_1, t_2, z)$  must have at least one maximum uniqueness of maximum can be tested as in the paper I and the part I of this paper.

3. *Influence of Variations of Environments on the System*

As before, in order to take into account of slow changes in environment, the equation of continuity (in average) for energy can be written as

$$d\bar{E} = \sum_i \frac{\partial \bar{E}}{\partial x_i} dx_i + d'Q, \quad \dots \dots \dots (33)$$

where  $x_i$ 's are the parameters involved in the definitions of the environments and

$$\left(-\frac{\partial \bar{E}}{\partial x_i}\right) = -\frac{1}{\left(\log \frac{1}{z_0}\right)} \frac{\partial}{\partial x_i} \{\log f(t_{10}, t_{20}, z_0)\}. \quad \dots \dots (34)$$

Then,

$$\begin{aligned} \left(\log \frac{1}{z_0}\right) d'Q &= \left(\log \frac{1}{z_0}\right) \left\{d\bar{E} - \sum \frac{\partial \bar{E}}{\partial x_i} dx_i\right\} \\ &= d \left\{\left(\log \frac{1}{z_0}\right) \bar{E}\right\} + \frac{\partial}{\partial z_0} \{\log f(t_{10}, t_{20}, z_0)\} + \sum_i \frac{\partial}{\partial x_i} \{\log f(t_{10}, t_{20}, z_0)\} \\ &= d \left\{\left(\log \frac{1}{z_0}\right) E_0 + \log f(t_{10}, t_{20}, z_0) - M_{10} \log t_{10} - M_{20} \log t_{20}\right\}. \dots (35) \end{aligned}$$

Now, according to the argument advanced in the paper I,

$$z_0 = e^{-\frac{1}{kT}}. \quad \dots \dots \dots (36)$$

Then, the entropy is given by

$$S = \frac{E}{kT} + \log F(t_{10}, t_{20}, T) - M_{10} \log t_{10} - M_2 \log t_{20}. \quad \dots \dots (37)$$

4. *Laws of Microscopic Distributions*

Now, as before, for deducing microscopic distributions, the system will be assumed to be composed of  $N_1$  and  $N_2$  numbers of particles, viz. molecules or atoms, where

$$M_1 = N_1 m_1, M_2 = N_2 m_2. \quad \dots \dots \dots (38)$$

If the number of ways, in which a particular  $n_1$ , (number of molecules of type (1)) is realized with energy  $\epsilon_{1p}$ , when the state of system is specified by  $M_{10}, M_{20}$

and  $E_0$ , is denoted by  $w'_1(E - n_{1\nu}\epsilon_{1\nu}, N_1 - n_{1\nu}, N_2)$ , then the expected value of  $n_{1\nu}$  is given by

$$\begin{aligned}
 n_{1\nu} &= \frac{\sum_{E=0}^{\infty} \sum_{N_1=0}^{\infty} \sum_{N_2=0}^{\infty} \sum_{n_{1\nu}=0}^{N_1} n_{1\nu} w'(E - n_{1\nu}\epsilon_{1\nu}, N_1 - n_{1\nu}, N_2) t_1'^{N_1} t_2'^{N_2} z_0^E}{\sum_{E=0}^{\infty} \sum_{N_1=0}^{\infty} \sum_{N_2=0}^{\infty} W(E, N_1, N_2) t_1'^{N_1} t_2'^{N_2} z_0^E} \\
 &= \frac{\left( \sum_{n_{1\nu}=0}^{\infty} n_{1\nu} t_1'^{n_{1\nu}} z_0^{n_{1\nu}\epsilon_{1\nu}} \right) \left[ 1 + \sum_E \sum_{N_1} \sum_{N_2} w'(E, N_1, N_2) t_1'^{N_1} t_2'^{N_2} z_0^E \right]}{\left( 1 + \sum_{n_{1\nu}=1}^{\infty} t_1'^{n_{1\nu}} z_0^{n_{1\nu}\epsilon_{1\nu}} \right) \left[ 1 + \sum_E \sum_{N_1} \sum_{N_2} w'(E, N_1, N_2) t_1'^{N_1} t_2'^{N_2} z_0^E \right]} \\
 &= \frac{t_1' \frac{\partial}{\partial t_1} \left( 1 + \sum_{n_{1\nu}=1}^{\infty} t_1'^{n_{1\nu}} z_0^{n_{1\nu}\epsilon_{1\nu}} \right)}{\left( 1 + \sum_{n_{1\nu}=0}^{\infty} t_1'^{n_{1\nu}} z_0^{n_{1\nu}\epsilon_{1\nu}} \right)} \\
 &= t_1' \frac{\partial}{\partial t_1} \left\{ \log \left( 1 + \sum_{n_{1\nu}=0}^{\infty} t_1'^{n_{1\nu}} z_0^{n_{1\nu}\epsilon_{1\nu}} \right) \right\} \\
 &= t_1' \frac{\partial}{\partial t_1} \left\{ \log g_\nu(t_1', z_0) \right\}, \quad \dots \quad \dots \quad \dots \quad \dots \quad \dots \quad \dots \quad \dots \quad (39)
 \end{aligned}$$

where

$$\left. \begin{aligned}
 t_1' &= t_{10}^{m_1} \\
 t_2' &= t_{20}^{m_2}
 \end{aligned} \right\} \dots \dots \dots (40)$$

Similar are the expressions for  $n_{2\nu}$ .

Thus, usual discussions for microscopic distributions will follow as before. The discussion on fluctuations can also be made and standard existing formulae (Tolman, 1946) can be deduced as in the paper I.

*General Concluding Remarks*

From the discussions in this paper, it is evident that the extension of the method proposed in the paper I to the general cases, viz. of the multi-component systems containing radiation in absence of chemical reactions, is easy and straightforward. Thus, this method, without losing any of the advantageous characteristics, mentioned in the paper I, is powerful enough to yield the usual formulae in the form as given by Fowler (1936) for general cases. The discussions of crystals and of chemical reactions by this method will be subject-matter of future communications.

**ABSTRACT**

In this paper, the thermodynamic behaviour of systems of different types, viz. those of black-body radiation, of matter and radiation, and of mixtures, has been investigated by the general statistical method, developed in a paper (Dutta, 1953) previously. Expressions for



thermodynamic functions, microscopic distributions, fluctuations, etc., have been obtained in the form given by Fowler. As in the previous paper (Dutta, 1953), it is seen that for investigations of microscopic properties, the introduction of any assumption about microscopic nature is not at all necessary. This assumption is to be introduced only for discussions of microscopic distributions or the like.

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