

# GEOMETRODYNAMICS OF THE NULL FIELD

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For a non-holonomic form of the matrix  $((h_{ij}))$ , compatible with the real electromagnetic null field, the conditions of existence of two analytic functions  $\rho$  and  $\sigma$  satisfying the Einstein-Maxwell equations are investigated.

§ 1. *Introduction*: In a four-dimensional space-time of general relativity we assume a symmetric tensor  $h_{\lambda\mu}$  of signature  $+++--$  and rank 4. There exists an inverse tensor  $h^{\lambda\mu}$  such that

$$h^{\lambda\mu}h_{\lambda\alpha} = \delta_{\alpha}^{\mu} \quad \dots \quad \dots \quad \dots \quad \dots \quad (1.1)$$

we shall use  $h_{\lambda\mu}$  ( $h^{\lambda\mu}$ ) to lower (raise) the indices. Throughout this paper the Latin indices  $i, j, k$  are used for the non-holonomic frame and range from 1 to 4 and the Greek indices  $\alpha, \beta, \gamma$  for the original frame and range from I to IV. The range of Greek indices  $\xi, \eta$ , is I, II and III.

If  $k_{\lambda\mu} = -k_{\mu\lambda}$  is a skew-symmetric real electromagnetic tensor field, then the stress tensor  $F_{\mu}^{\lambda}$  is given by

$$F_{\mu}^{\lambda} \stackrel{\text{def}}{=} k_{\mu}^{\alpha}k_{\alpha}^{\lambda} - \frac{1}{4}\delta_{\mu}^{\lambda}k_{\alpha\beta}k^{\beta\alpha}. \quad \dots \quad \dots \quad \dots \quad (1.2)$$

We put

$${}^{(p)}k_{\mu}^{\alpha} \stackrel{\text{def}}{=} k_{\mu}^{\beta}k_{\beta}^{\alpha}, \quad {}^{(0)}k_{\mu}^{\alpha} \stackrel{\text{def}}{=} \delta_{\mu}^{\alpha} \quad (p = 1, 2, \dots) \quad \dots \quad \dots \quad (1.3a)$$

$$4K \stackrel{\text{def}}{=} {}^{(2)}k_{\alpha}^{\alpha}. \quad \dots \quad \dots \quad \dots \quad \dots \quad \dots \quad (1.3b)$$

Then the basic system of field equations of general relativity assumes the form

$$R_{\mu}^{\lambda} - \frac{1}{2}Rh_{\mu}^{\lambda} = \mu u_{\mu}u^{\lambda} - F_{\mu}^{\lambda} + p\delta_{\mu}^{\lambda} \quad \dots \quad \dots \quad \dots \quad (1.4a)$$

and the Maxwell equations are

$$\nabla_{[\omega}k_{\mu\lambda]} = 0 \quad \dots \quad \dots \quad \dots \quad \dots \quad (1.4b)$$

$$\nabla_{\alpha}k^{\alpha\beta} = \epsilon u^{\beta} \quad \dots \quad \dots \quad \dots \quad \dots \quad (1.4c)$$

where  $\nabla_{\alpha}$  is the covariant derivative with respect to  $h_{\lambda\mu}$

$$R \stackrel{\text{def}}{=} R_{\alpha}^{\alpha} \quad \dots \quad \dots \quad \dots \quad \dots \quad (1.5)$$

and  $u^{\alpha}$  is a unit time-like vector field

$$U_{\alpha}U^{\alpha} = -1 \quad \dots \quad \dots \quad \dots \quad \dots \quad (1.6)$$

and  $p$  is the pressure and  $\epsilon$  is the energy.

The electromagnetic tensor  $k_{\lambda\mu}$  is said to be of the

- (a) first class if  $k = \det |k_{\alpha\beta}| \neq 0, K \neq 0$   
 (b) second class if  $k = 0, K \neq 0$   
 (c) third class if  $k = 0, K = 0$

If  $\lambda$  be an eigen-value of  $k_{\alpha}^{\beta}$  and  $a_{\alpha}$  be the corresponding eigen-vector, then

$$|k_{\alpha}^{\beta} - \lambda \delta_{\alpha}^{\beta}| a_{\beta} = 0 \quad \dots \quad \dots \quad \dots \quad \dots \quad (1.7)$$

Thus the eigen-values are given by

$$|k_{\alpha}^{\beta} - \lambda \delta_{\alpha}^{\beta}| = 0 \quad \dots \quad \dots \quad \dots \quad \dots \quad (1.8)$$

which reduces to

$$\lambda^4 + 2K\lambda^2 + k = 0. \quad \dots \quad \dots \quad \dots \quad \dots \quad (1.9)$$

Hlavaty (1961) has shown that when  $k_{\alpha\beta}$  is of the first and second class a set of four mutually independent eigen-vectors  $a_{\beta}$  can always be found, which, together with its inverse set  $a^{\beta}$ , form a frame which is called the non-holonomic frame of the first two classes. For the third class which includes the null electromagnetic field, there is only one basic scalar  $\lambda = 0$ , and therefore a non-holonomic frame cannot be found out of the basic vectors. However, Hlavaty (1961) using line geometry has shown that for the third class also a set of two complex conjugate null vectors  $a_1^{\lambda}, a_2^{\lambda}$  and two real null vectors  $a_3^{\lambda}, a_4^{\lambda}$ , all linearly independent can be found such that

$$\left. \begin{aligned} a_1^{\lambda} k_{\lambda}^{\mu} &= \frac{1}{\sqrt{2}} a^{\mu} e^{+i\sigma} \\ a_3^{\lambda} k_{\lambda}^{\mu} &= \frac{1}{\sqrt{2}} (e^{+i\sigma} a_1^{\mu} + e^{+i\sigma} a_2^{\mu}) \\ a_4^{\lambda} k_{\lambda}^{\mu} &= 0 \end{aligned} \right\} \dots \quad \dots \quad \dots \quad (1.10)$$

Mishra (1963) has obtained this result without using line geometry, with the help of the property

$${}^{(3)}k_{\alpha}^{\beta} = 0.$$

The non-holonomic form for the canonical matrix  $((h_{ij}))$  compatible with real magnetic null field was given by Mishra (1963) as

$$\begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 0 & -1 \\ 0 & 0 & -1 & 0 \end{pmatrix} \quad \dots \quad \dots \quad \dots \quad \dots \quad (1.11)$$

§ 2. *Determination of  $k_{\lambda\mu}$  and  $u_\lambda$* : We take a generalization of form (1.1) as

$$((h_{ij})) = \begin{pmatrix} \alpha & 0 & 0 & \rho \\ 0 & \beta & 0 & 0 \\ 0 & 0 & 0 & -\alpha \\ \sigma & 0 & -\alpha & \sigma \end{pmatrix} \quad \dots \quad \dots \quad \dots \quad (2.1)$$

$$\det ((h_{\lambda\mu})) = -\alpha^3\beta$$

$$h_{\mu\nu} \stackrel{\text{def}}{=} h_{ij} a_\mu^i a_\nu^j \quad \dots \quad \dots \quad \dots \quad \dots \quad (2.2)$$

where  $\alpha, \beta, \rho, \sigma$  are four parameters of which  $\alpha$  and  $\beta$  are to be determined by Maxwell equations. We then have

$$((h^{ij})) = \begin{pmatrix} \frac{1}{\alpha} & 0 & \frac{\rho}{\alpha^2} & 0 \\ 0 & \frac{1}{\beta} & 0 & 0 \\ \frac{\rho}{\alpha^2} & 0 & \frac{\alpha\rho^2 - \sigma}{\alpha^3} & -\frac{1}{\alpha} \\ 0 & 0 & -\frac{1}{\alpha} & 0 \end{pmatrix} \quad \dots \quad \dots \quad \dots \quad (2.3)$$

Since the field is null and  $k_{ij}$  is skew-symmetric, using conditions (Mishra 1963)

$$k_1^4 k_4^2 = k_1^2 k_2^4 = k_4^1 k_2^4 = k_2^1 k_1^4 = k_4^2 k_2^1 = k_4^1 k_1^2 = 0 \quad \dots \quad (2.4a)$$

and

$$k_4^1 k_1^3 + k_4^2 k_2^3 = 1. \quad \dots \quad \dots \quad \dots \quad (2.4b)$$

We obtain for the non-holonomic frame (2.1)

$$k_1^3 = k_4^1 = k_4^2 = \alpha, \quad k_2^3 = \beta \quad \dots \quad \dots \quad \dots \quad (2.5a)$$

$$k_3^2 = k_1^4 = k_3^4 = k_3^1 = k_2^4 = k_2^1 = k_1^2 = 0 \quad \dots \quad \dots \quad (2.5b)$$

$$\alpha(\alpha + \beta) = 1 \quad \dots \quad \dots \quad \dots \quad \dots \quad (2.5c)$$

We get

$$((k_{ij})) = \begin{pmatrix} 0 & 0 & 0 & -\alpha^2 \\ 0 & 0 & 0 & -\alpha\beta \\ 0 & 0 & 0 & 0 \\ \alpha^2 & \alpha\beta & 0 & 0 \end{pmatrix} \quad \dots \quad \dots \quad \dots \quad (2.6)$$

consequently

$$\begin{aligned} (2)k_{\lambda\mu} &\stackrel{\text{def}}{=} (2)k_{ij} a_\lambda^i a_\mu^j \\ &= -\alpha a_\lambda^4 a_\mu^4 \quad \dots \quad \dots \quad \dots \quad (2.7) \end{aligned}$$

Since

$$a_\lambda^i = h_{ij} a_\lambda^j$$

we have

$$\left. \begin{aligned} a_\lambda^1 &= \alpha a_\lambda + \rho a_\lambda^4 \\ a_\lambda^2 &= \beta^2 a_\lambda \\ a_\lambda^3 &= -\alpha a_\lambda \\ a_\lambda^4 &= \rho a_\lambda - \alpha a_\lambda + \sigma a_\lambda^4 \end{aligned} \right\} \dots \dots \dots (2.8)$$

and

$$\left. \begin{aligned} a_\lambda^1 &= \frac{1}{\alpha} a_\lambda + \frac{\rho}{\alpha} a_\lambda^4 \\ a_\lambda^2 &= \frac{1}{\beta} a_\lambda \\ a_\lambda^3 &= \frac{\rho}{\alpha} a_\lambda + \frac{\rho^2 - \sigma}{\alpha^2} a_\lambda^4 - \frac{1}{\alpha} a_\lambda^4 \\ a_\lambda^4 &= -\frac{1}{\alpha} a_\lambda \end{aligned} \right\} \dots \dots \dots (2.9)$$

From (2.8) and (2.9) we get,

$$h^{ij} = h^{\lambda\mu} a_\lambda^i a_\mu^j \dots \dots \dots (2.10)$$

From (2.2)

$$h_{\mu\nu} = \alpha a_\mu^1 a_\nu^1 + 2\rho a_\mu^1 a_\nu^4 + \beta a_\mu^2 a_\nu^2 + 2\alpha a_\mu^3 a_\nu^4 + \sigma a_\mu^4 a_\nu^4 \dots \dots (2.11)$$

Using (2.7), (1.4a) can now be written as

$$R_{\mu\lambda} - (\frac{1}{2}R + p)h_{\mu\lambda} = \mu U_\mu U_\lambda - \alpha a_\mu^4 a_\lambda^4 \dots \dots (2.12)$$

The right-hand side of (2.12) is of rank two, therefore the left-hand side must be of second rank. Therefore we can take

$$U_\lambda = A a_\lambda^3 + B a_\lambda^4 \dots \dots \dots (2.13)$$

( $a_\lambda^1$  and  $a_\lambda^2$  are excluded since they are complex conjugate) where  $A$  and  $B$  have to be determined. We have

$$U_\lambda U^\lambda = -1 = \frac{A(\rho^2 - \alpha\sigma)}{\alpha^3} - \frac{2AB}{\alpha} \dots \dots (2.14)$$

If we put  $AB = 1$ , then

$$U_\lambda = A a_\lambda^3 + \frac{1}{A} a_\lambda^4 \dots \dots \dots (2.15)$$

where

$$A = \pm \alpha \sqrt{\frac{2 - \alpha}{\rho^2 - \alpha\sigma}}$$

From the Maxwell equations (1.4b) and (1.4c) we have

$$\nabla_{[i} k_{j]k} = 0 \dots \dots \dots (2.16)$$

$$\nabla_i k_i^i = 0 \dots \dots \dots (2.17)$$

Substituting from (2.6) in (2.16) we get

$$\nabla_3 \alpha = 0 \quad \dots \dots \dots (2.18a)$$

$$\nabla_3 \beta = 0 \quad \dots \dots \dots (2.18b)$$

Since  $\alpha$  and  $\beta$  are scalars it follows that  $\alpha$  and  $\beta$  are constant along the  $X^3$  axis. From (2.17), substituting from (2.5), we get

$$\left. \begin{aligned} \nabla_3 \alpha &= \epsilon u_1 = 0 \\ \nabla_3 \beta &= \epsilon u_2 = 0 \\ \nabla_1 \alpha &= \epsilon u_4 \neq 0 \end{aligned} \right\} \dots \dots \dots (2.19)$$

Therefore, if there is energy, then the first and second components of  $u_j$  are zero while the fourth component is non-zero.

§ 3. Using (2.7) and (2.11) eqn. (1.4) can be written as

$$\begin{aligned} R_{\mu\nu} - \left(\frac{R}{2} + p\right) (\alpha a_{\mu}^1 a_{\nu}^1 + 2\rho a_{\mu}^1 a_{\nu}^2 + \beta a_{\mu}^2 a_{\nu}^2 + 2\alpha a_{\mu}^4 a_{\nu}^1) \\ = \left(\frac{\mu}{A^2} - \sigma - \alpha\right) a_{\mu}^4 a_{\nu}^4 + \mu A^2 a_{\mu}^3 a_{\nu}^3 + \mu (a_{\nu}^3 a_{\mu}^4 + a_{\nu}^4 a_{\mu}^3) \end{aligned} \dots (3.1)$$

In order to simplify eqn. (3.1) we write

$$h_{ij} = \tilde{h}_{ij} + X_{ij} \quad \dots \dots \dots (3.2)$$

where

$$((\tilde{h}_{ij})) = \begin{pmatrix} \alpha & 0 & 0 & 0 \\ 0 & \beta & 0 & 0 \\ 0 & 0 & 0 & -\alpha \\ 0 & 0 & -\alpha & 0 \end{pmatrix} \quad \dots \dots \dots (3.3)$$

$$\det |\tilde{h}_{ij}| = \alpha^3 \beta$$

and

$$((X_{ij})) = \begin{pmatrix} 0 & 0 & 0 & \rho \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ \rho & 0 & 0 & \sigma \end{pmatrix} \quad \dots \dots \dots (3.4)$$

$$\tilde{h}_{\lambda\mu} \stackrel{\text{def}}{=} \tilde{h}_{ij} a_{\lambda}^i a_{\mu}^j \quad \dots \dots \dots (3.5a)$$

$$X_{\lambda\mu} \stackrel{\text{def}}{=} X_{ij} a_{\lambda}^i a_{\mu}^j \quad \dots \dots \dots (3.5b)$$

If  $\Gamma_{\lambda\mu}^{\nu}$  are the connections formed with  $h_{\mu\nu}$  and  $\tilde{\Gamma}_{\lambda\mu}^{\nu}$  are the connections formed with  $\tilde{h}_{\mu\nu}$  and  $\tilde{\nabla}_{\mu}$  stands for the covariant derivative with respect to  $\tilde{h}_{\mu\nu}$  then we make use of the following results (Schmidt 1966):

$$T_{\lambda\mu|\nu} \stackrel{\text{def}}{=} \frac{1}{2} (\tilde{\nabla}_{\lambda} T_{\mu\nu} + \tilde{\nabla}_{\mu} T_{\lambda\nu} - \tilde{\nabla}_{\nu} T_{\lambda\mu}) \quad \dots \dots \dots (3.6a)$$

for any tensor  $T_{\lambda\mu}$

$$T_{\lambda\mu}{}^\nu \stackrel{\text{def}}{=} T_{\lambda\mu|\alpha} h^{\alpha\nu} \quad \dots \quad \dots \quad \dots \quad \dots \quad (3.6b)$$

$$h_{\lambda\mu}{}^\nu = X_{\lambda\mu}{}^\nu \quad \dots \quad \dots \quad \dots \quad \dots \quad (3.6c)$$

$$\Gamma_{\lambda\mu}^\nu = \tilde{\Gamma}_{\lambda\mu}^\nu + X_{\lambda\mu}{}^\nu \quad \dots \quad \dots \quad \dots \quad \dots \quad (3.6d)$$

$$X_{\lambda\mu}{}^\nu = 0 \quad \dots \quad \dots \quad \dots \quad \dots \quad (3.6e)$$

The equations (3.1) become on using (3.6)

$$\begin{aligned} \tilde{R}_{\mu\nu} - \tilde{\nabla}_\alpha X_{\mu\nu}{}^\alpha + X_{\alpha\mu}{}^\beta - \left(\frac{R}{2} + p\right) (\tilde{h}_{\mu\nu} + X_{\mu\nu}) \\ = \left(\frac{\mu}{A^2} - \alpha\right) a_\mu a_\nu + \mu A^2 a_\mu a_\nu + \mu (a_\nu a_\mu + a_\nu a_\mu) \end{aligned} \quad \dots \quad (3.7)$$

§ 4. In order to apply the Cauchy existence theorem we separate in (3.7) the second derivatives from the rest (Schmidt 1966). We get

$$h^{\alpha\beta} (\partial_{\mu\alpha} X_{\nu\beta} + \partial_{\nu\alpha} X_{\mu\beta} - \partial_{\alpha\beta} X_{\mu\nu} - \partial_{\mu\nu} X_{\alpha\beta}) = Q_{\mu\nu} \quad \dots \quad (4.1)$$

where  $Q_{\mu\nu}$  denote the rest of the equations. Now separating the derivatives  $\partial_{\text{IV IV}}$  from the rest and letting the indices  $\xi, \eta$  run through I, II and IH only, we get

$$h^{\text{IV IV}} \partial_{\text{IV IV}} X_{\xi\eta} = M_{\xi\eta} \quad \text{for } \mu, \nu = \xi, \eta \quad \dots \quad (4.2)$$

$$h^{\text{IV } \eta} \partial_{\text{IV IV}} X_{\xi\eta} = M_{\xi \text{IV}} \quad \text{for } \mu, \nu = \text{IV}, \xi \quad \dots \quad (4.3)$$

$$h^{\xi\eta} \partial_{\text{IV IV}} X_{\xi\eta} = M_{\text{IV IV}} \quad \text{for } \mu, \nu = \text{IV}, \text{IV} \quad \dots \quad (4.4)$$

where  $M_{\mu\nu}$  stands for the rest.

From (3.4) and (3.5b) the above equations become on substitution

$$(\partial_{\text{IV IV}} \sigma) a_\xi a_\eta + 2(\partial_{\text{IV IV}} \rho) a_\xi a_\eta = N_{\xi\eta} \quad \dots \quad (4.5)$$

$$h^{\text{IV } \eta} [(\partial_{\text{IV III}} \sigma) a_\xi a_\eta + 2(\partial_{\text{IV IV}} \rho) a_\xi a_\eta] = N_{\text{IV } \xi} \quad \dots \quad (4.6)$$

$$h^{\xi\eta} [(\partial_{\text{IV IV}} \sigma) a_\xi a_\eta + 2(\partial_{\text{IV IV}} \rho) a_\xi a_\eta] = N_{\text{IV IV}} \quad \dots \quad (4.7)$$

where  $N_{\mu\nu}$  stands for the rest.

Now if we denote the dual vectors to  $a_\xi$  by  $a^\xi$  then

$$a_\xi a^\xi = \delta_\xi^\xi$$

$$a_\xi a^\eta = \delta_\xi^\eta \quad \dots \quad \dots \quad \dots \quad (4.8)$$

where  $r = 1, 2, 3$ .

Equations (4.5), (4.6) and (4.7) become

$$\left. \begin{aligned} \partial_{IV IV} \sigma &= a^{\xi} a^{\eta} N_{\xi \eta} \\ \partial_{IV IV} \rho &= a^{\xi} a^{\eta} N_{\xi \eta} \end{aligned} \right\} \dots \dots \dots (4.9)$$

and

$$\left. \begin{aligned} N_{\xi \eta} a^{\xi} a^{\eta} &= 0 \\ h^{IV \eta} N_{\xi \eta} &= N_{\xi IV} \\ h^{\xi \eta} N_{\xi \eta} &= N_{IV IV} \end{aligned} \right\} \dots \dots \dots (4.10)$$

for  $(r, s) = (1, 1), (1, 2), (1, 3), (2, 2), (2, 3)$  and  $(3, 3)$ .

A unique solution  $(\rho, \sigma)$  can be obtained by prescribing boundary values for  $\rho, \sigma, \partial_{VI} \rho, \partial_{IV} \sigma$  along the hypersurface  $X^{IV} = 0$  such that the boundary values satisfy (4.10) (Schmidt 1966).

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