

LITHIUM IN LUNA-24 SAMPLES

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Lithium contents for three Luna-24 samples are found to vary from 8 ppm to 10 ppm. Examination of Li values from various lunar missions show that more the KREEP content in a soil, the higher the Li value.

INTRODUCTION

Li abundances in lunar samples have been found to vary from 2 ppm to 100 ppm. However, in general, the Li values for lunar samples are higher than those of chondrites. It has been pointed out by some earlier works (Wänke *et al.*, 1977; Dreibus *et al.*, 1976, 1977) that Li does not condense with refractory elements but follows Mg, Fe silicates. Thermal volatilisation studies (Gibson & Hubbard, 1972) on lunar basalts, breccias and soil samples for alkali, alkaline and rare earth elements (REE) have shown that Li, Ba, Sr and REE do not volatilize below 1400 °C.

Several authors, in particular Wänke and co-workers (Wänke *et al.*, 1976 and earlier papers; Dreibus *et al.*, 1976, 1977) have discussed the Li abundances in detail. Earlier studies of Li in lunar samples have been described elsewhere (Shukla & Goel, 1974; Shukla, 1977). In the present paper, the present authors describe their results on three Luna-24 samples 24123, 24148, and 24190 allotted to us.

EXPERIMENTAL

Li contents for these soils have been measured by neutron activation using the reaction $\text{Li}^6 (n, \alpha)\text{H}^3$. Details of the experimental procedure are described elsewhere (Shukla *et al.*, 1978).

RESULTS AND DISCUSSIONS

Li results for three Luna-24 soils are given in Table I. Lithium values reported for other five Luna 24 soils by Nyquist *et al.* (1977) range from 5.6 ppm to 6.2 ppm and appear to be systematically low. It can be seen that Li values do not show any depth dependence. Earlier analyses of Luna-16 and Luna-20 samples exhibited the same trend e.g., for a depth variation of 0 to 33 cm, the Li values are practically constant at about 11 ± 1 ppm (Shukla & Goel, 1974). This observation alongwith the grain size independence of Li (Shukla, 1977) and no marked enrichment or depletion of Li in soils over rocks and breccias indicate that the various regolith processes do not seem to cause any redistribution of Li.

From earlier studies in this laboratory and elsewhere (Shukla, 1977; Wänke and co-workers) it has been observed that Li shows a positive correlation (in varying

TABLE I
Li values of Luna 24 samples*

Sample No.	Depth cm	Weight mg	Li ppm
24123-13	122-123	5.5	08.4
24148-11	147-148	8.0	10.0
24190-14	189-190	7.5	08.8
*Error \pm 10%			

TABLE II
Li values (ppm) for various lunar mission (soil samples)

Mission	Literature		This Laboratory	
	Range	Average	Range	Average
Apollo	11	—	11 ¹	—
	12	15-24	18 ²	—
	14	19-25	23 ³	—
	15	8-16	11 ⁴	16-17 (3 soils)
	16	6-10	8 ⁴	8-10
	17	7-11	9 ⁴	9-13
	17	7-11	9 ⁴	9-13
Luna	16	—	11 ⁵	10-12
	20	—	6 ⁶ , 6 ⁷	7 (2 soils)
	24	5.6-6.2	6 ⁸	8-10 (3 soils)

1. Annel and Helz (1970). 2. Cuttitta *et al.* (1971). 3. Rose *et al.* (1972). 4. Cuttitta *et al.* (1973) and Rose *et al.* (1973, 1974, 1975). 5. Vinogradov (1973). 6. Nava and Philpotts (1973). 7. Reed and Jovanovic (1973). 8. Nyquist *et al.* (1977).

degree) with elements like Be, Na, K, U, La, Zr and Nb. Of these elements, Li correlates strongly with Be and Nb. Due to lack of trace elements data for the soils studied in this work, various elemental correlations for these soils could not be checked.

In Table II, Li values for various Apollo and Luna missions are summarized. The highest Li values are observed for Apollo-14 samples which are mostly KREEP rich. Further, it may be pointed out that in Apollo-12 sample 12013, (pointed out to represent a KREEP sample by Hubbard *et al.*, 1971) highest values for Li (~100 ppm) have been observed for some density fractions. Thus it appears that the more the KREEP component in a soil, the higher the Li value. A low value of Li for the Luna-24 soils could be taken to indicate a low KREEP component for these soils.

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